



Uncovering Peculiar Double-Detonation Type Ia Supernovae

The rare weirdos: SN 2020jgb & SN 2022joj

Liu et al., 2023a, ApJ, 946, 83

Liu et al., 2023b, ApJ, 958, 178

Chang Liu, Adam Miller

2024.2



Type Ia Supernovae (SNe Ia)

- **Rare:** extragalactic
- **Bright:** luminosity $\sim$ the entire galaxy
- **Transient:** last for months



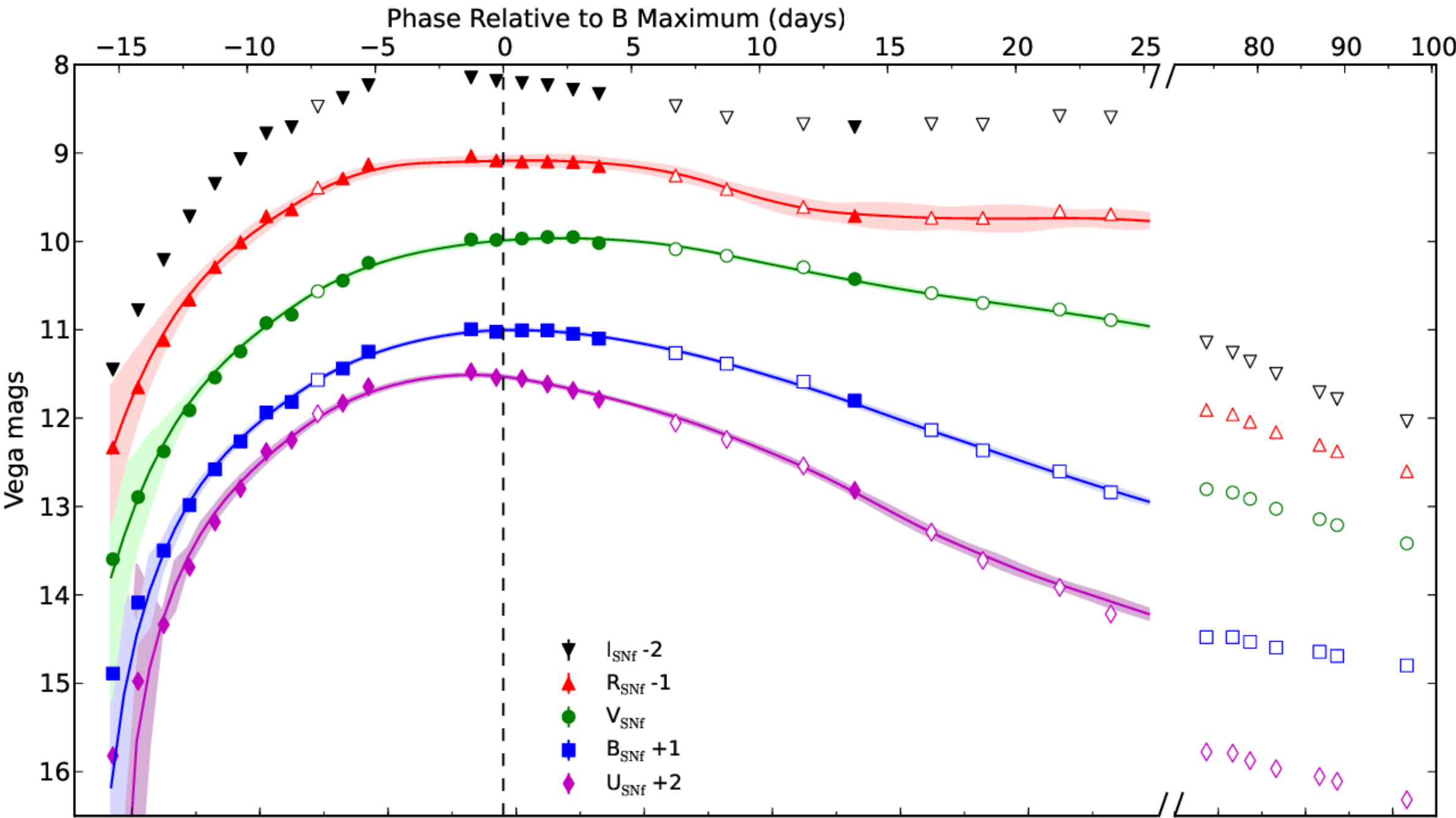
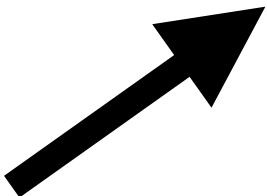
Zwicky Transient Facility

- Wide field camera
- Repeatedly observe the Northern sky every ~2 nights
- Detected, classified, and monitored $\sim 10^4$ SNe Ia since 2018

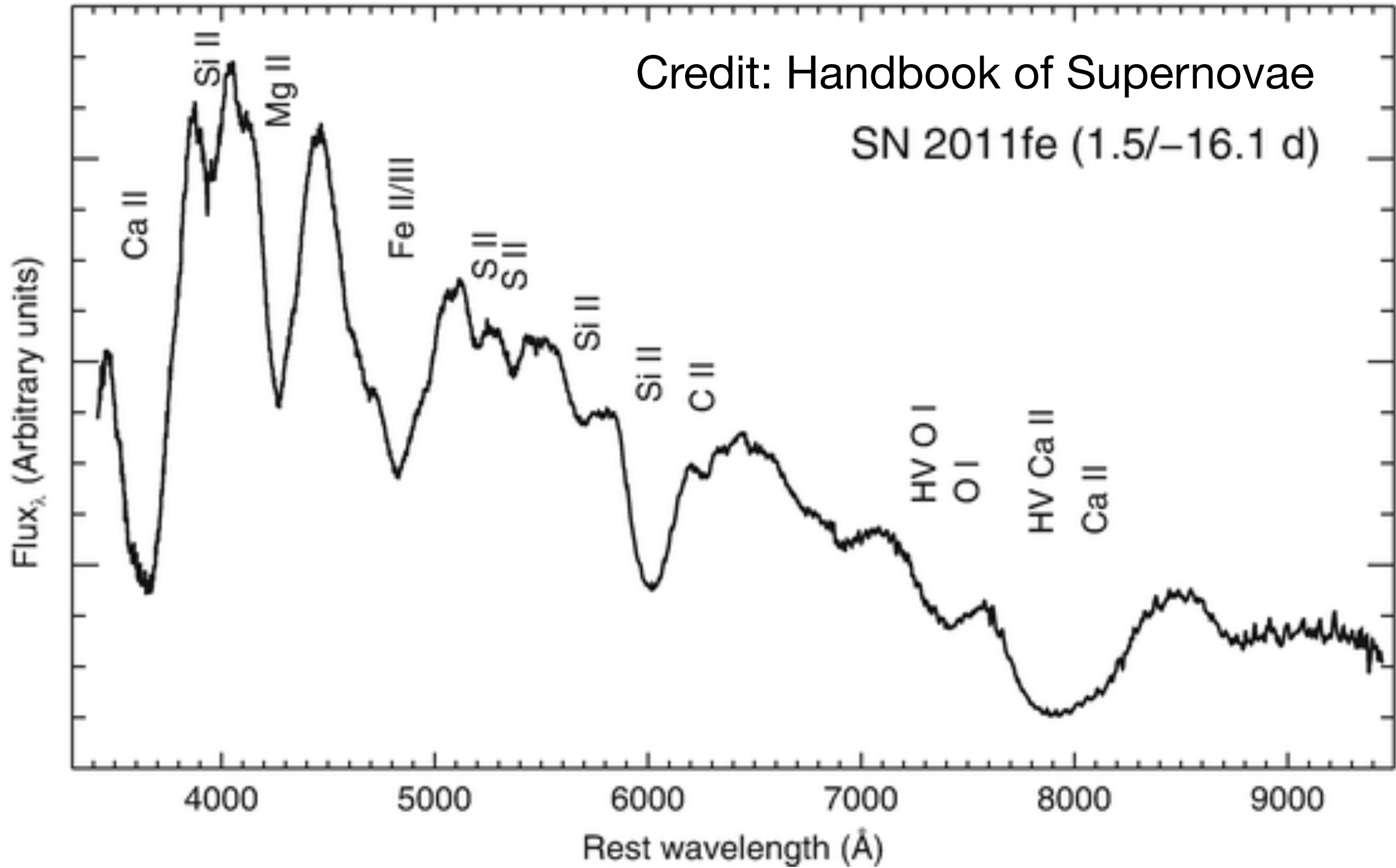
Image Credit: Palomar Observatory/Caltech

Type Ia Supernovae

Monitor cosmic explosions



Credit: Nearby Supernova Factory



An Unfinished Story...

How to explode a C/O white dwarf (WD)?

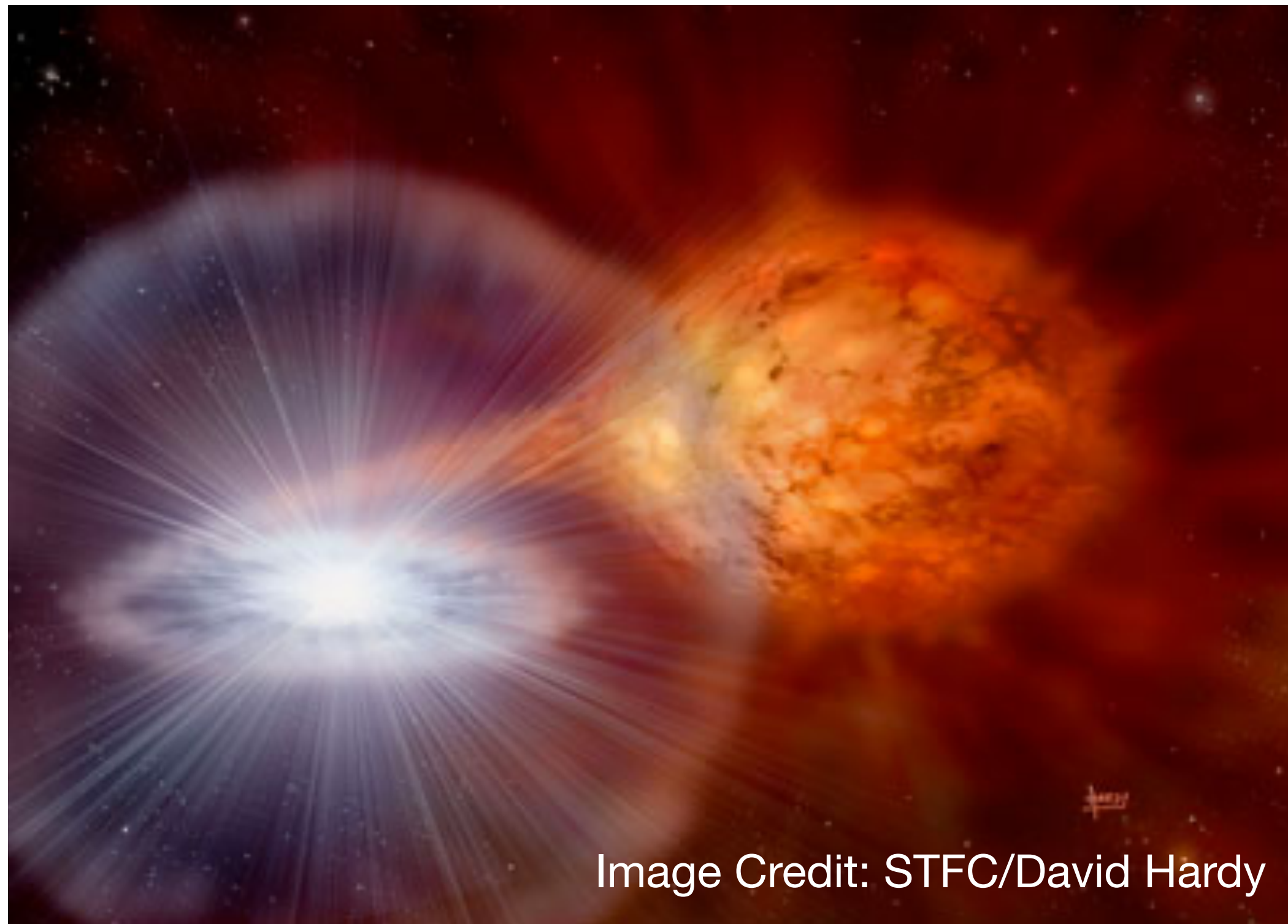
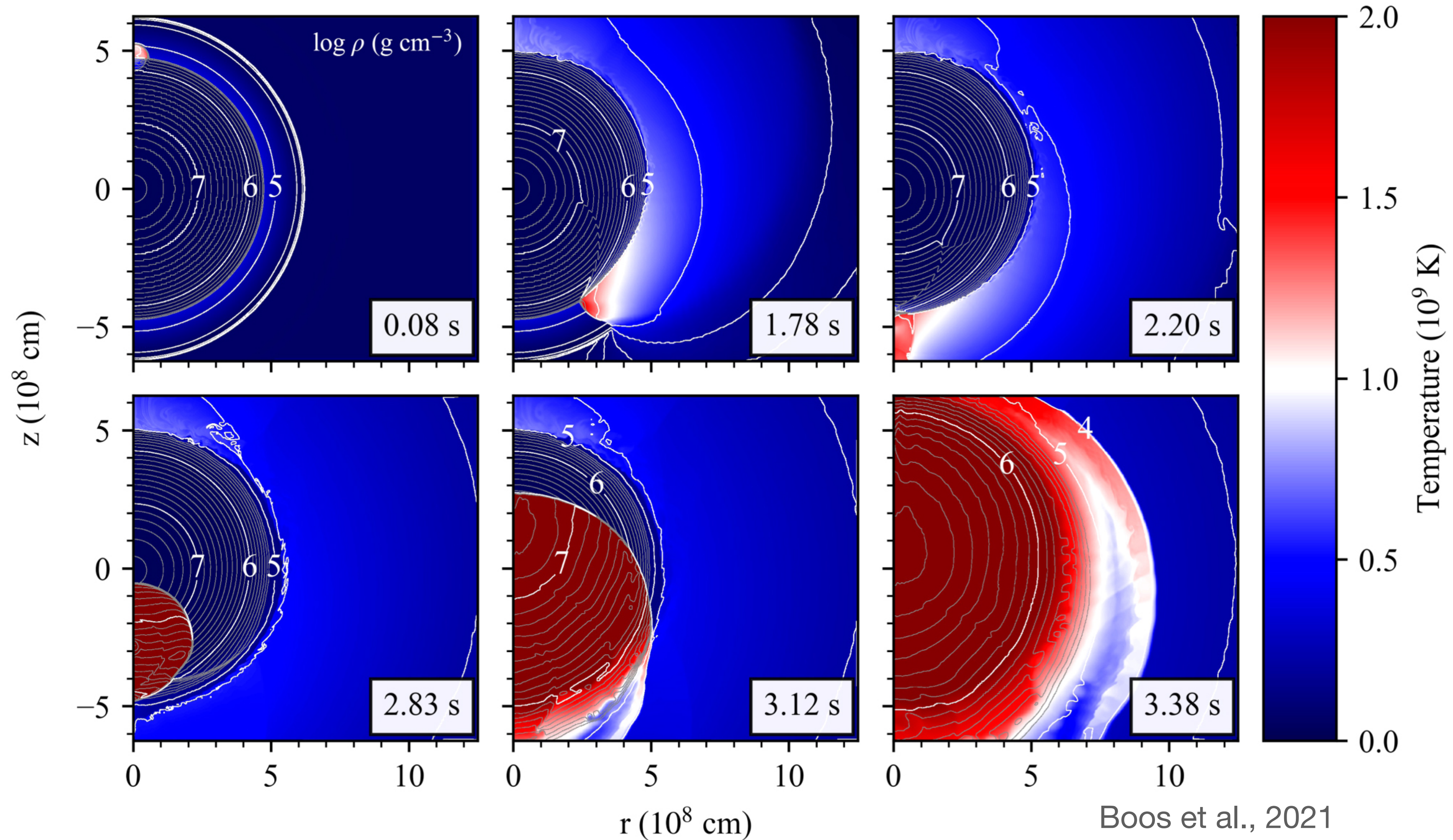


Image Credit: STFC/David Hardy

- Feeding the WD until it approaches **Chandrasekhar mass** ($M_{\text{Ch}} \sim 1.4 M_{\odot}$)
 - > Properties around maximum light ✓
 - > Event rates ✗
 - > Diversity in the population ✗
- Dynamical ignition in a **sub- M_{Ch}** WD

He-shell Double Detonation

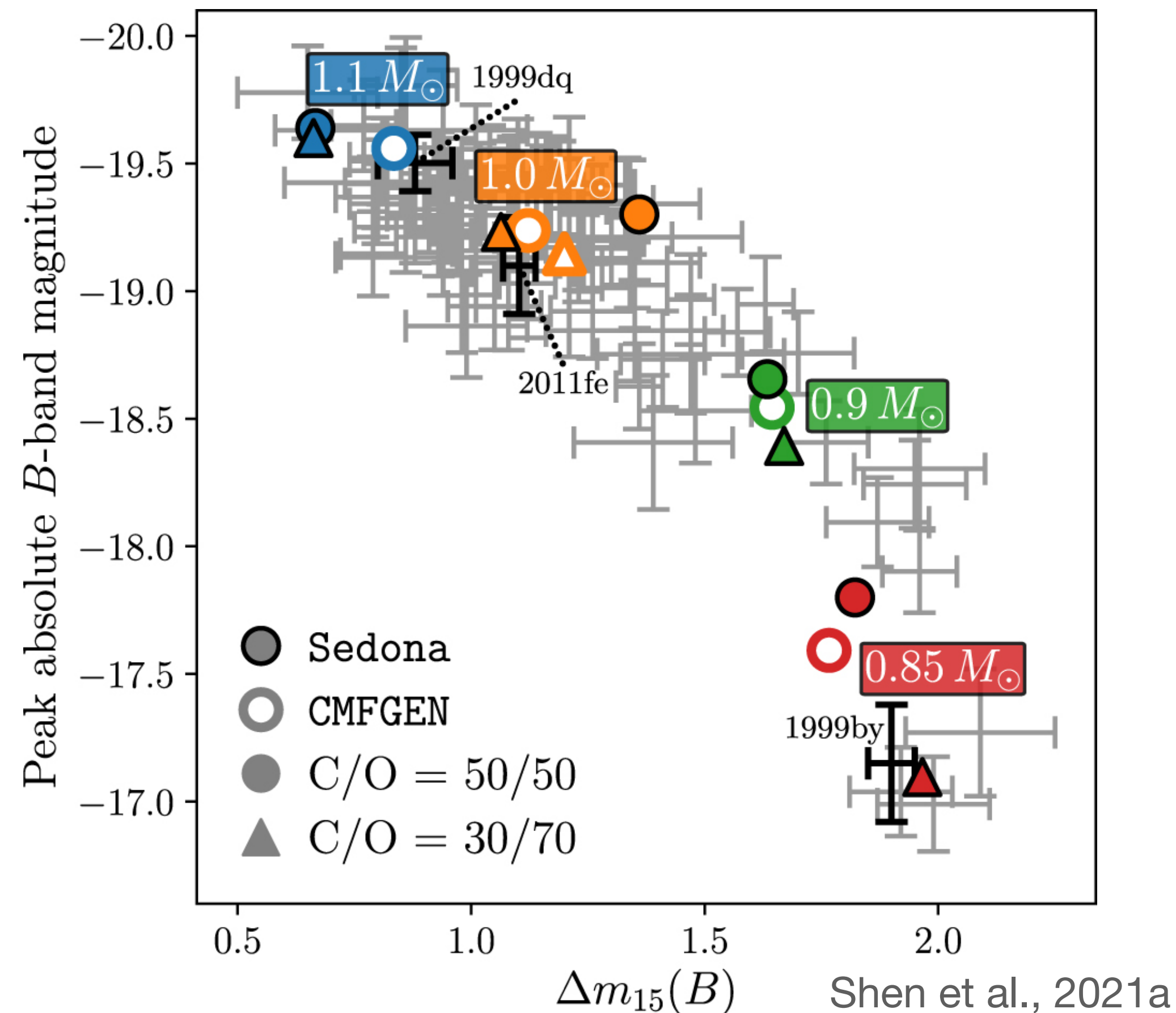
- Detonation of a helium shell triggers the **detonation** in the sub- M_{Ch} C/O WD



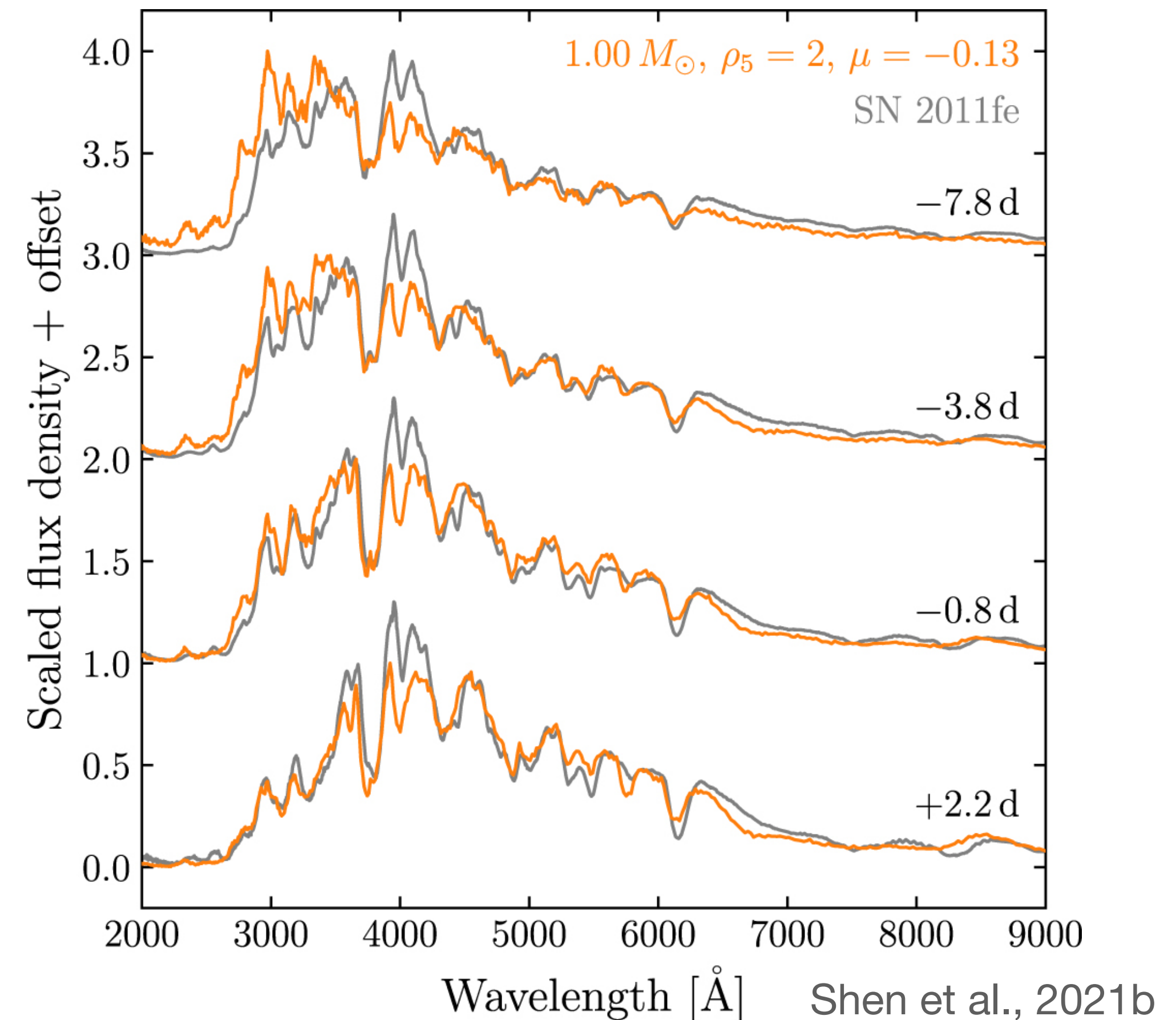
Thin He shell + sub- M_{ch} C/O core

Normal SNe Ia at peak? — under debate...

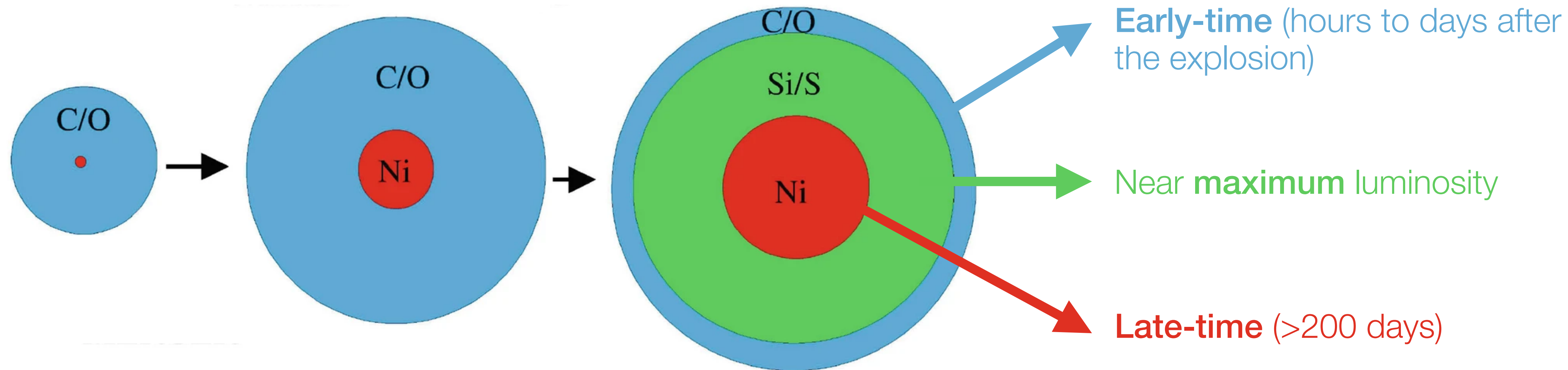
Luminosity-width (Phillips) relation



Spectroscopic features



Near- M_{Ch} WD



Sub- M_{Ch} WD in a double detonation

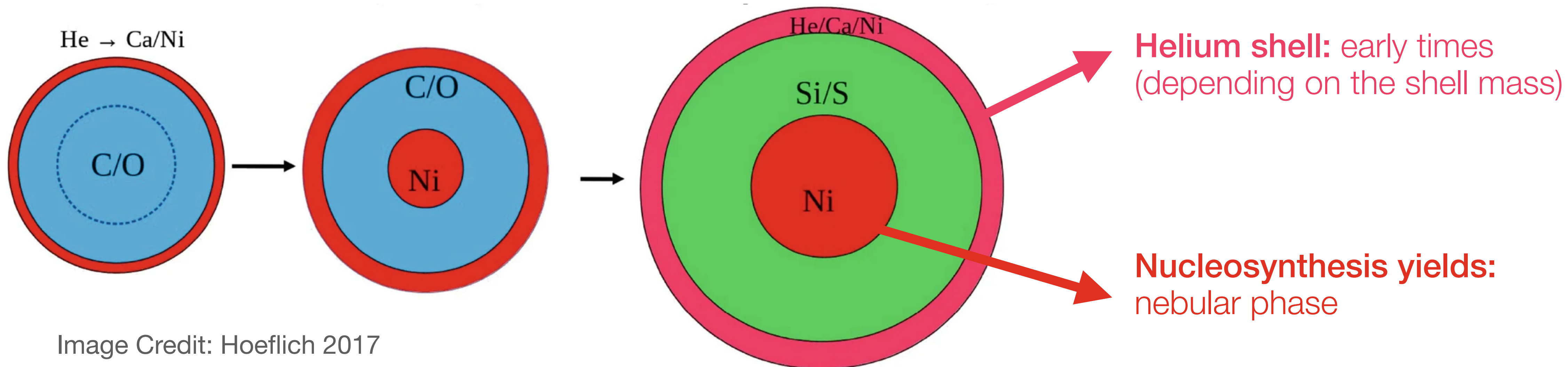


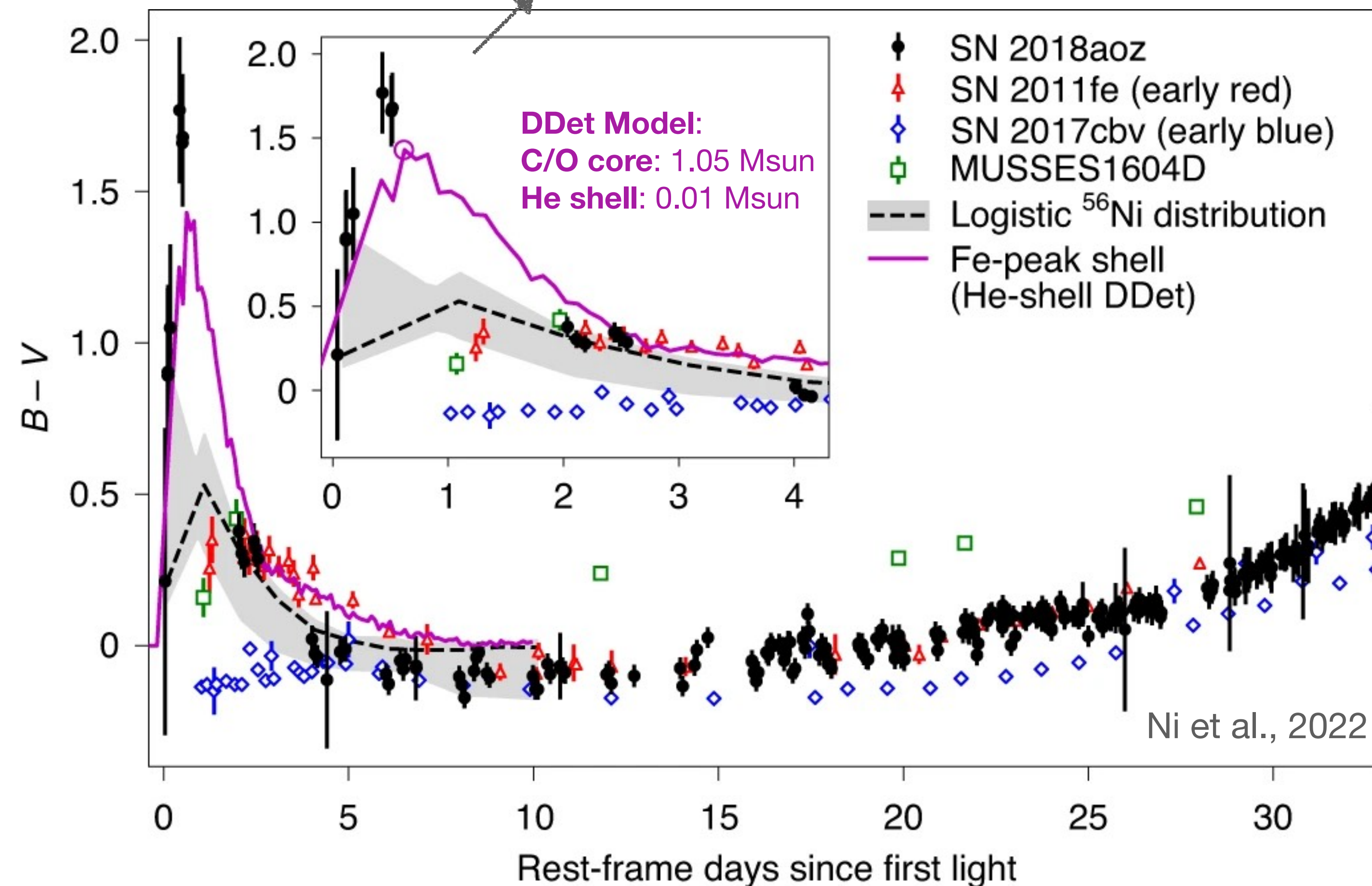
Image Credit: Hoeflich 2017

Thin He shell + sub- M_{ch} C/O core

Exotic early colors?

Fe-group elements
absorb UV photons

SN 2018aoz: Extremely red B-V color ~12 hr after the explosion

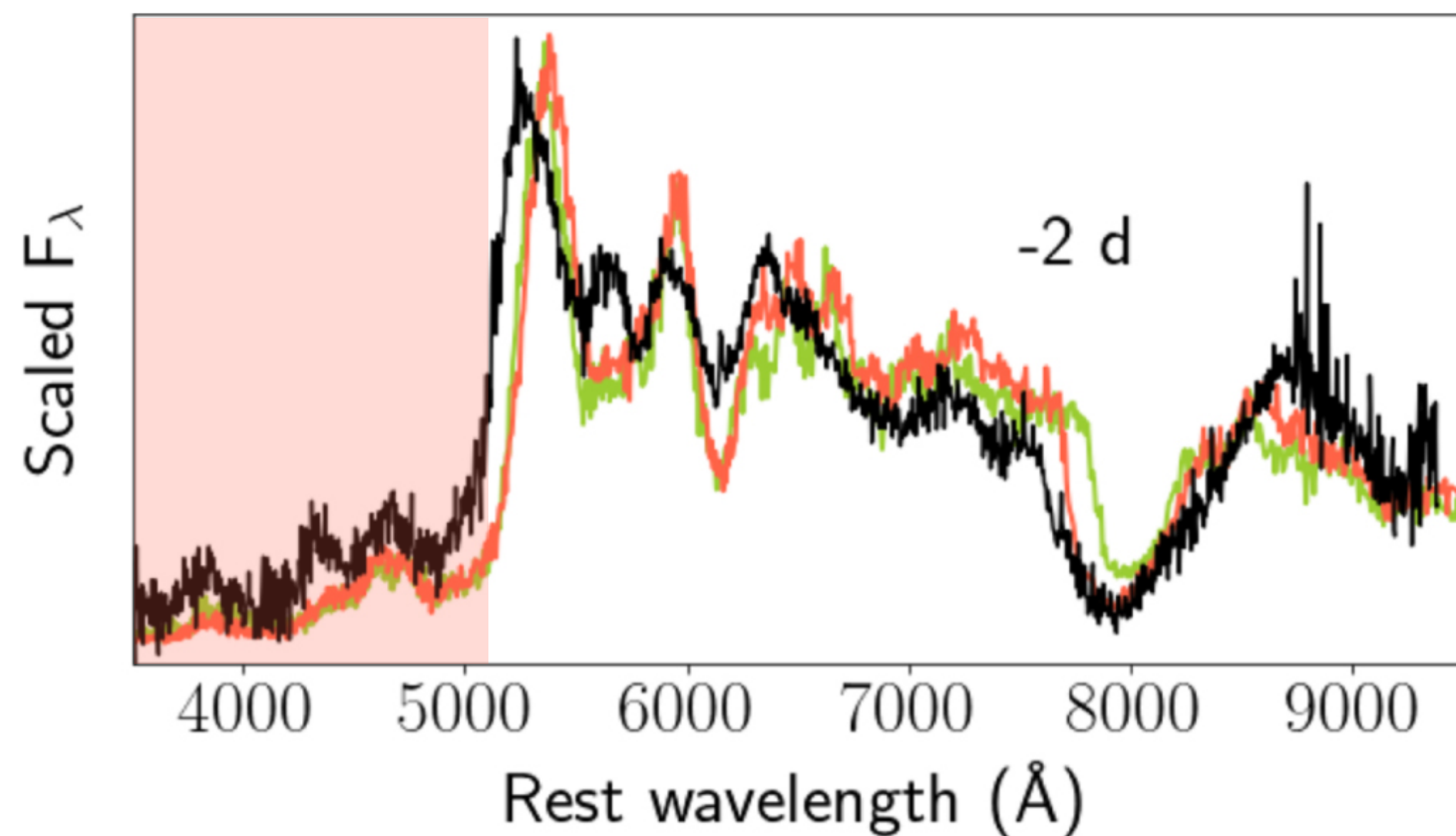


Massive He shell + sub- M_{ch} C/O core

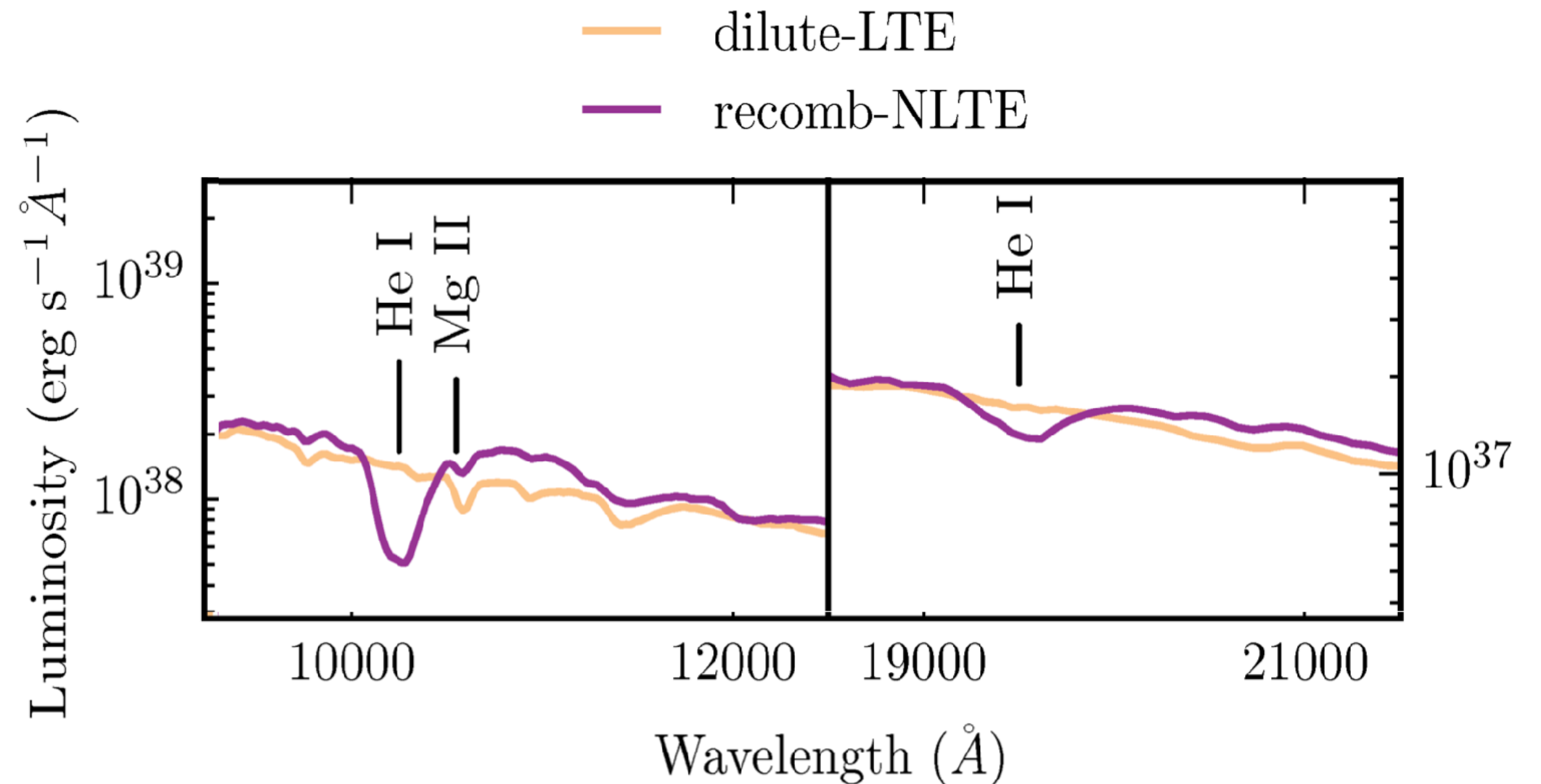
Peculiarities to expect **around maximum light**

Fe-group elements ← **He shell** → Unburnt helium

SN 2018byg: prototype



De et al., 2019



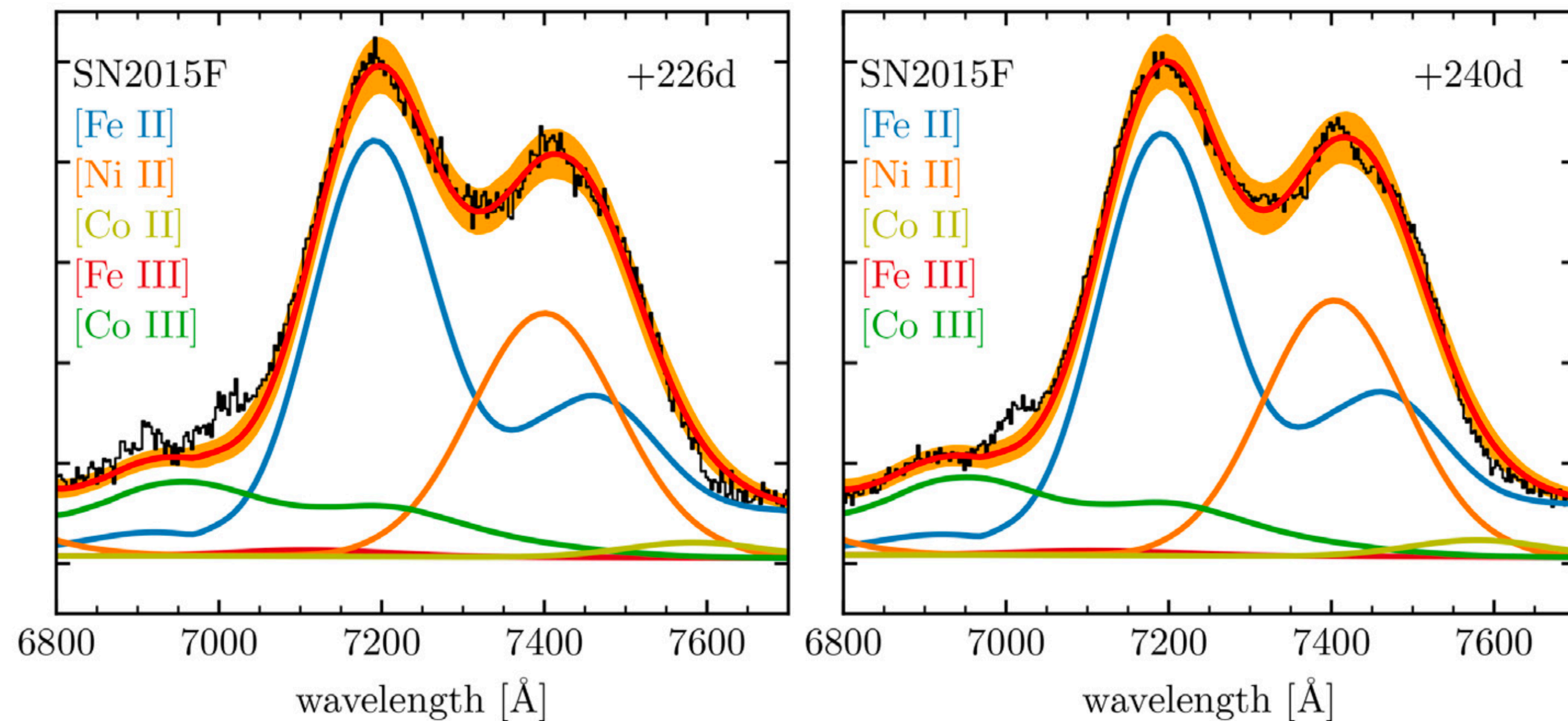
Boyle et al., 2017

Massive He shell + sub- M_{ch} C/O core

Peculiarities to expect **~ 1 yr after explosion**

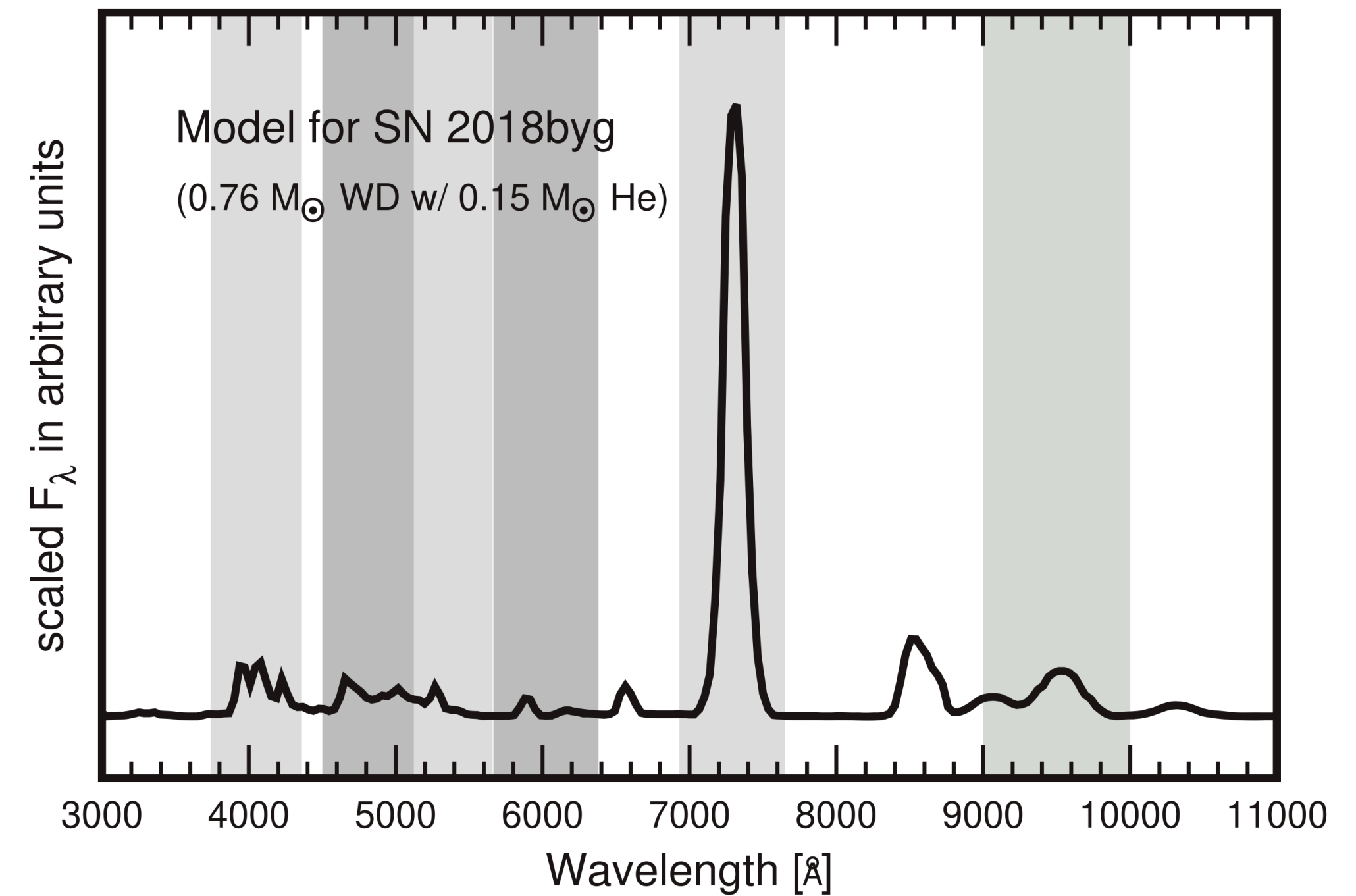
→ Ni-58/Ni-56

Low Ni/Fe abundance



Flörs et al., 2020

Strong [Ca II] emission



Polin et al., 2021

Massive He shell + sub- M_{ch} C/O core

From the **first** to the **last** photon

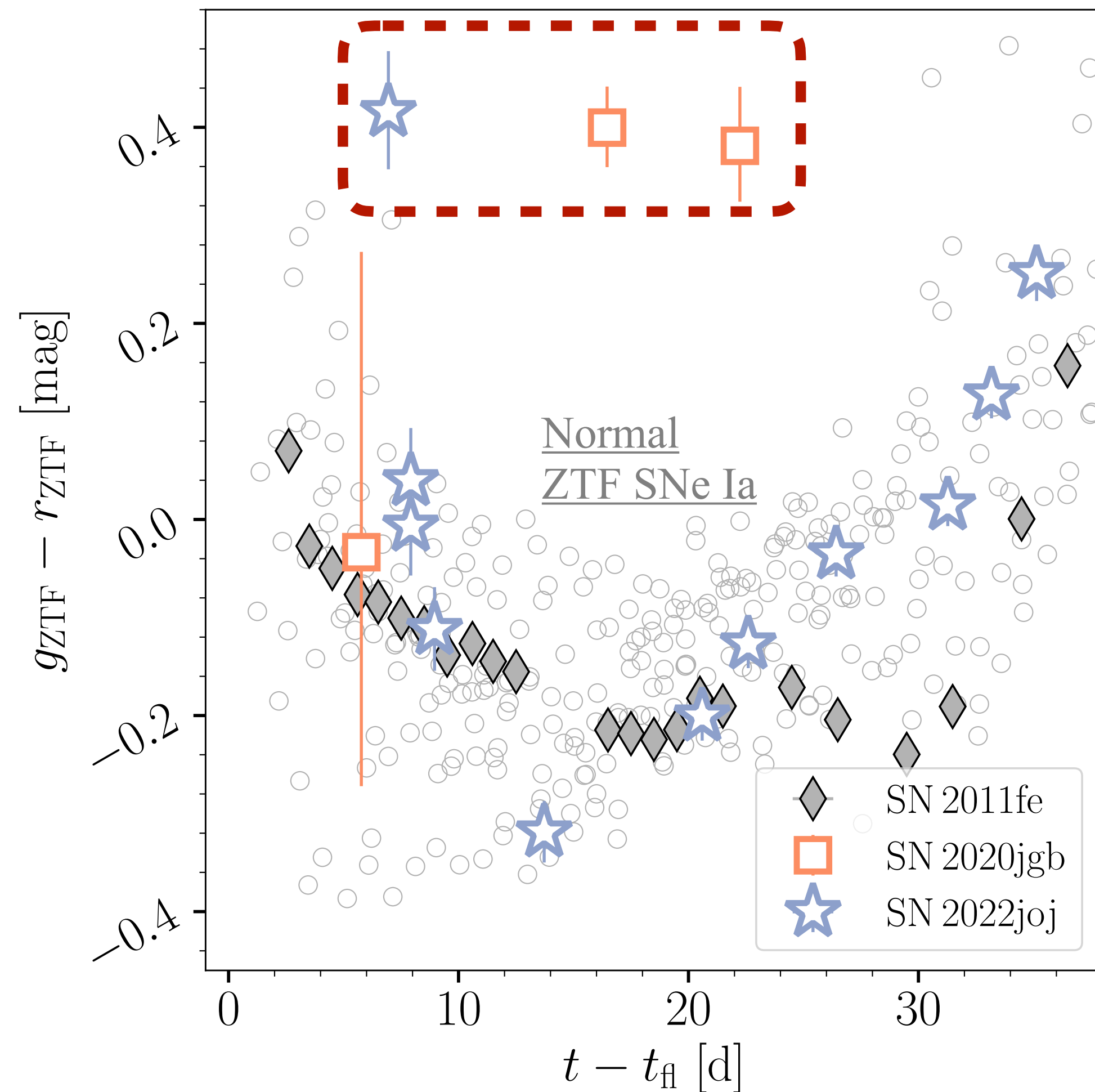
- **Days after explosion**
 - Flux excess — early discovery & photometry
- **Around maximum luminosity**
 - Dramatic color evolution — multi-band photometry
 - Line blanketing — optical spectroscopy
 - Unburnt helium — NIR spectroscopy
- **Late time**
 - [Ca II] emission, low Fe/Ni abundance ratio — deep spectroscopy

SNe 2020jgb & 2022joj - new members of this rare subclass discovered by ZTF!

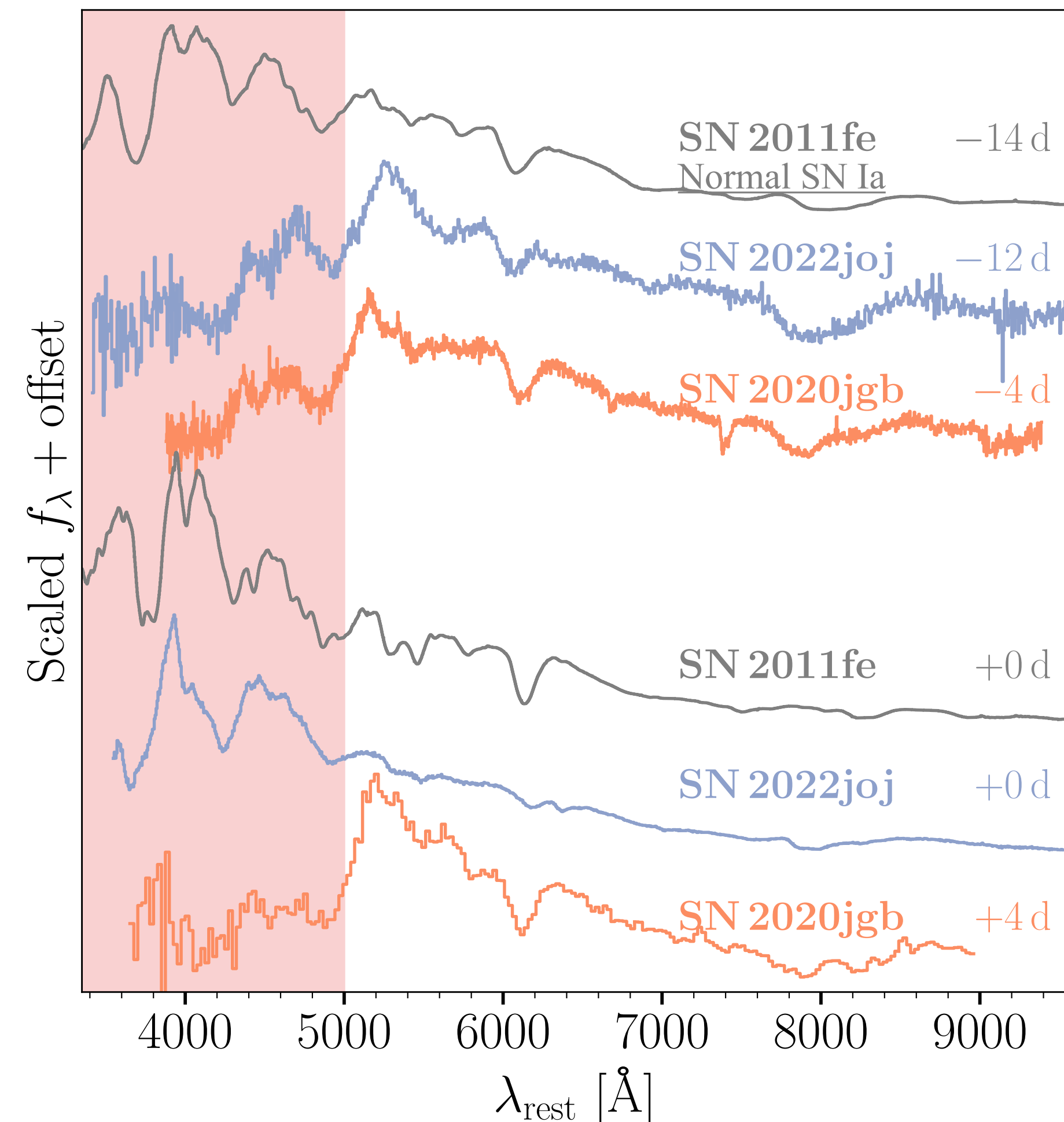
Fe-group Elements in the Shell

Line blanketing

Unusual **red** colors



Suppression of flux $< 5000 \text{ \AA}$

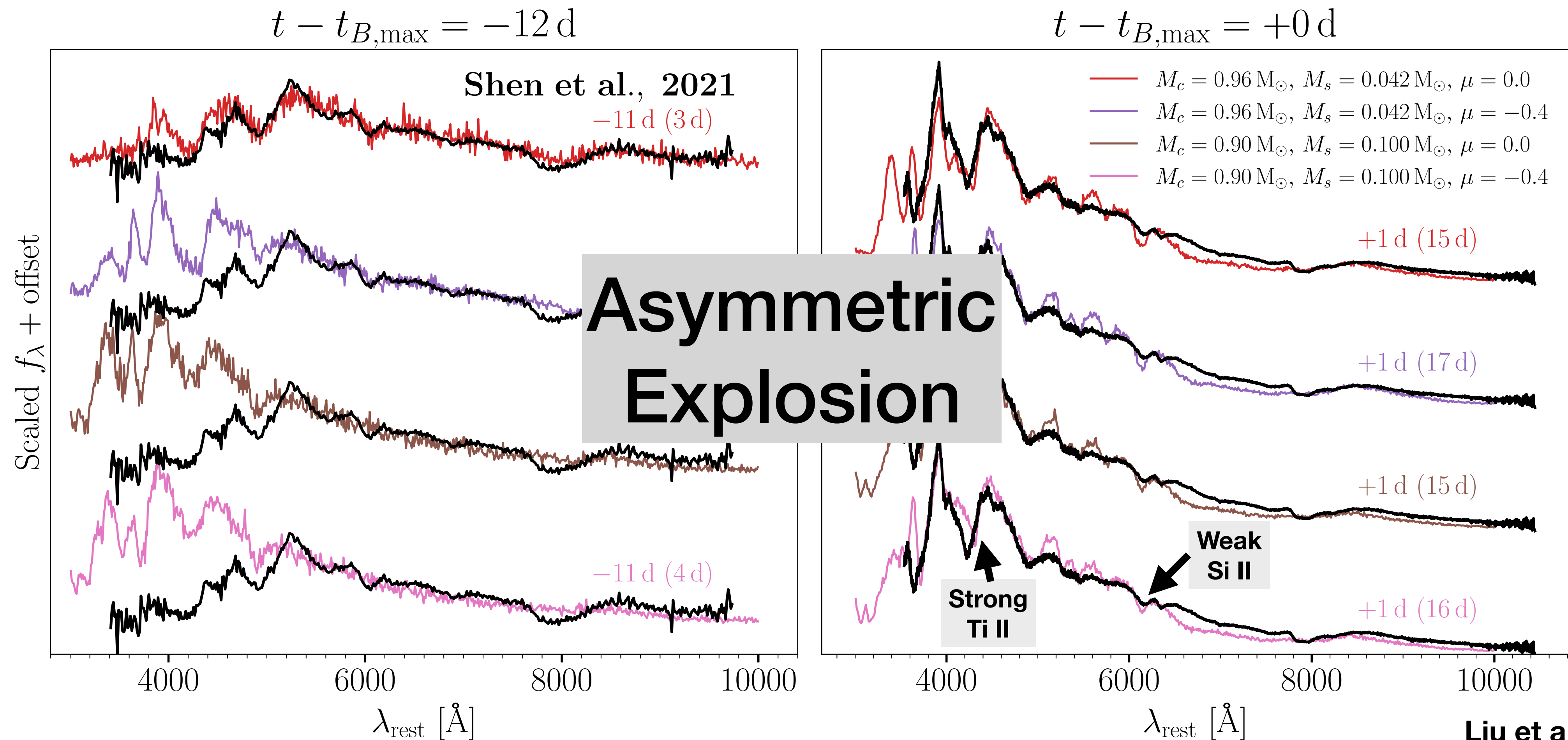


2D Simulations

Spectral modeling

SN 2022joj

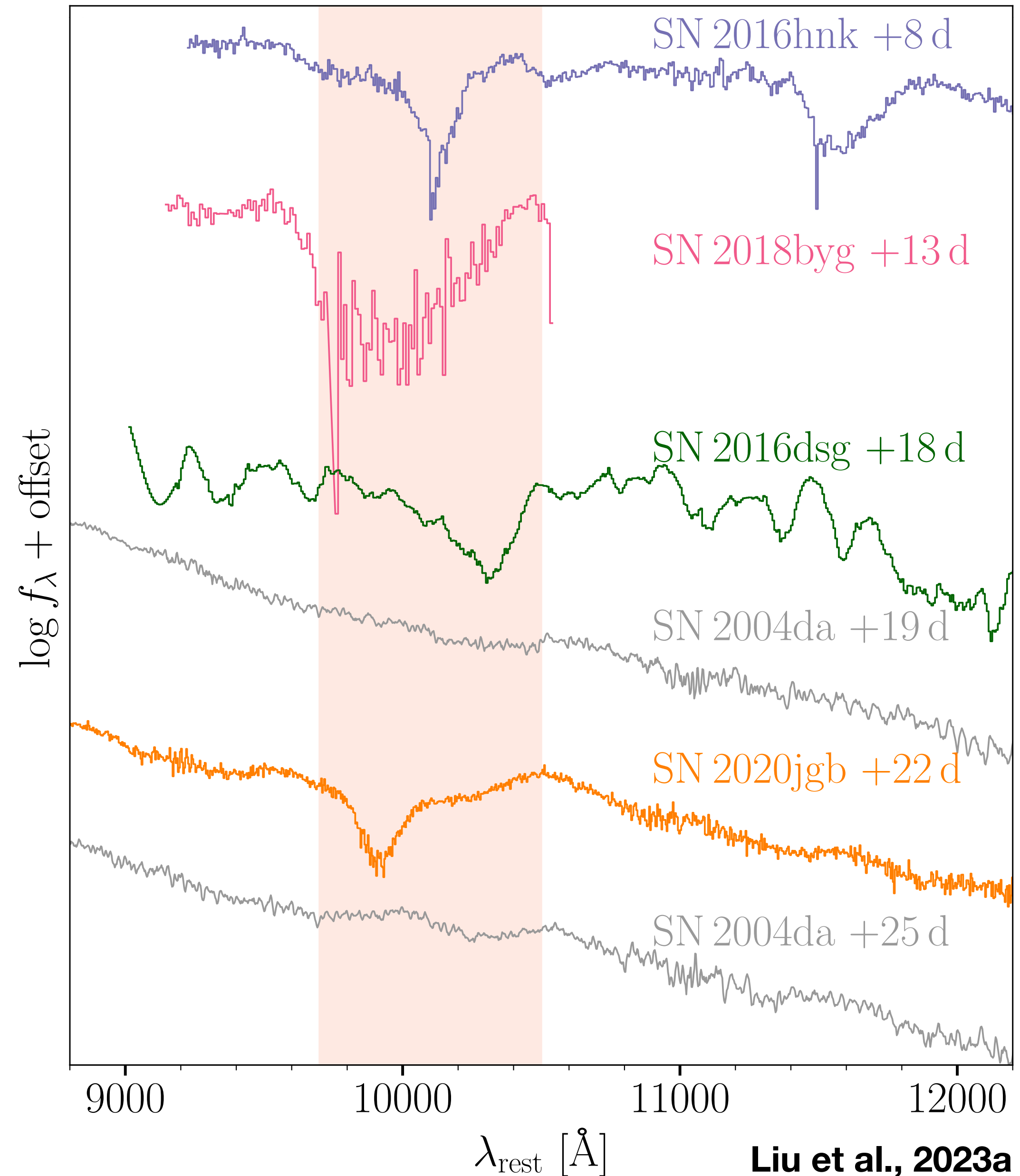
- $\sim 0.96 M_{\odot}$ (**core**) + $\sim 0.04 M_{\odot}$ (**shell**)
- Viewed from the **opposite** hemisphere of the shell detonation point



Unburnt Helium

He I $\lambda 10830$

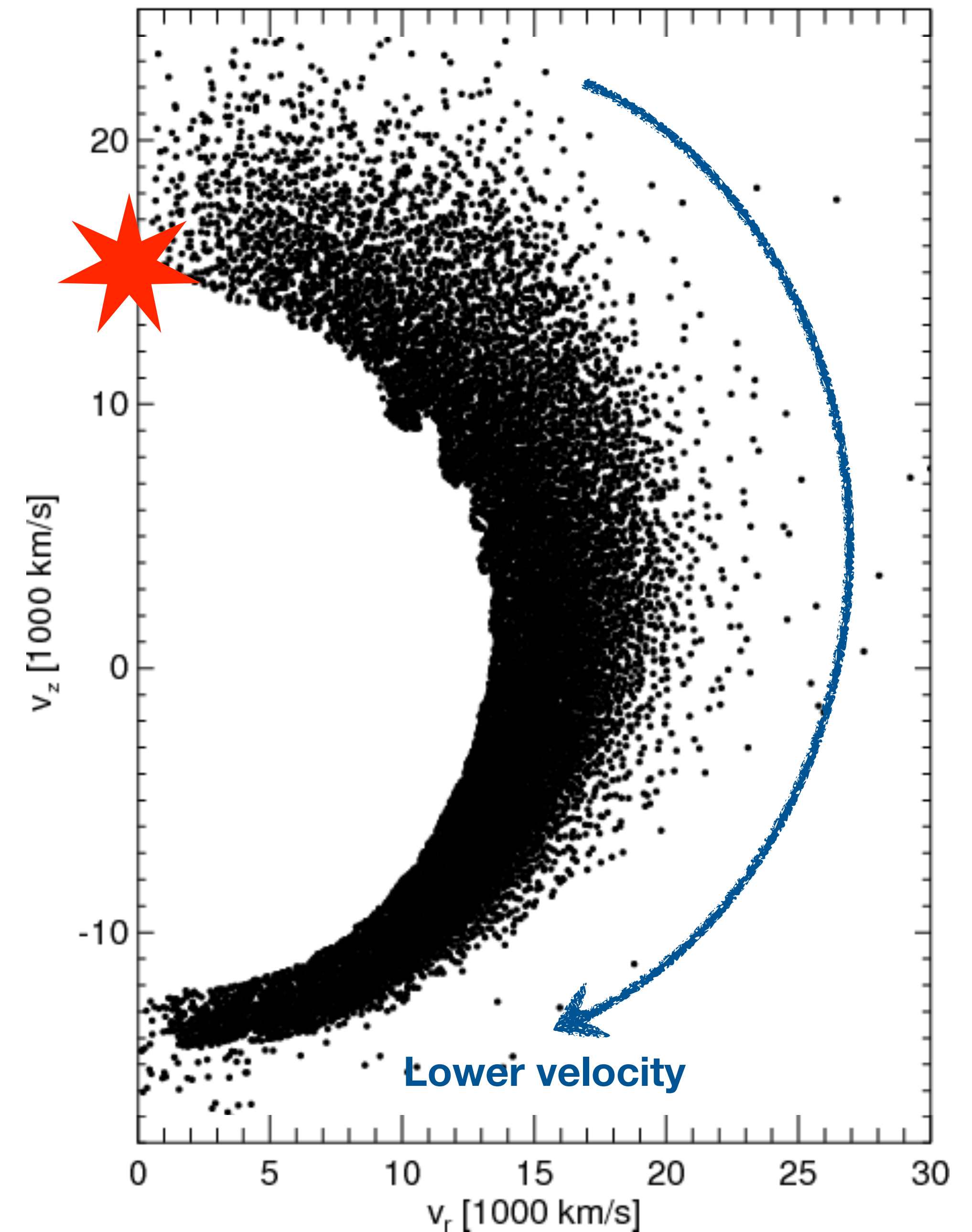
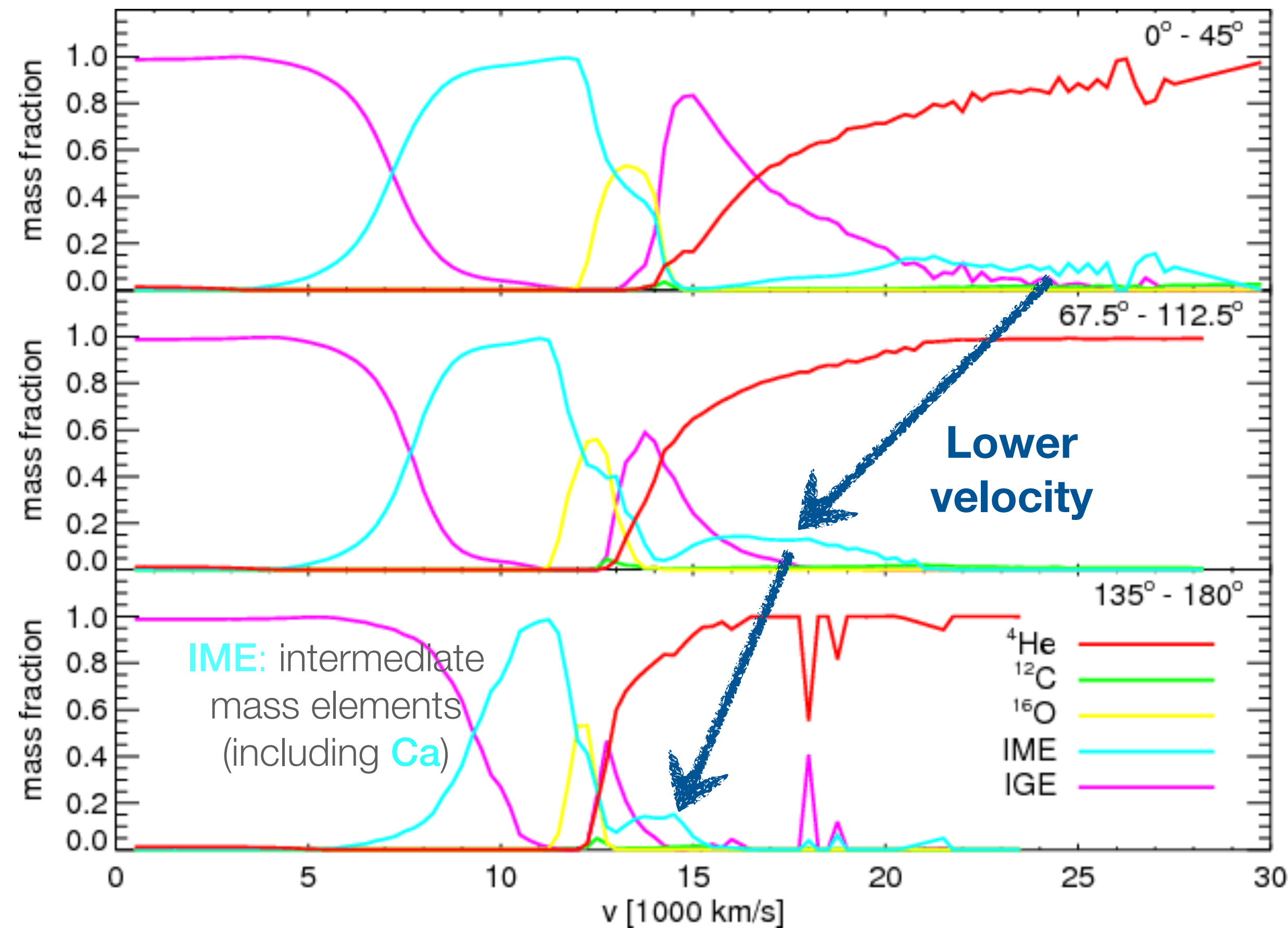
- Archival peculiar double detonation SNe: **<10**
- **Ubiquitous (?)** 1 micron features
 - **Helium** at a huge variety of **velocities?**
- Viewing angle effects?



Unburnt helium

Viewing angle effects

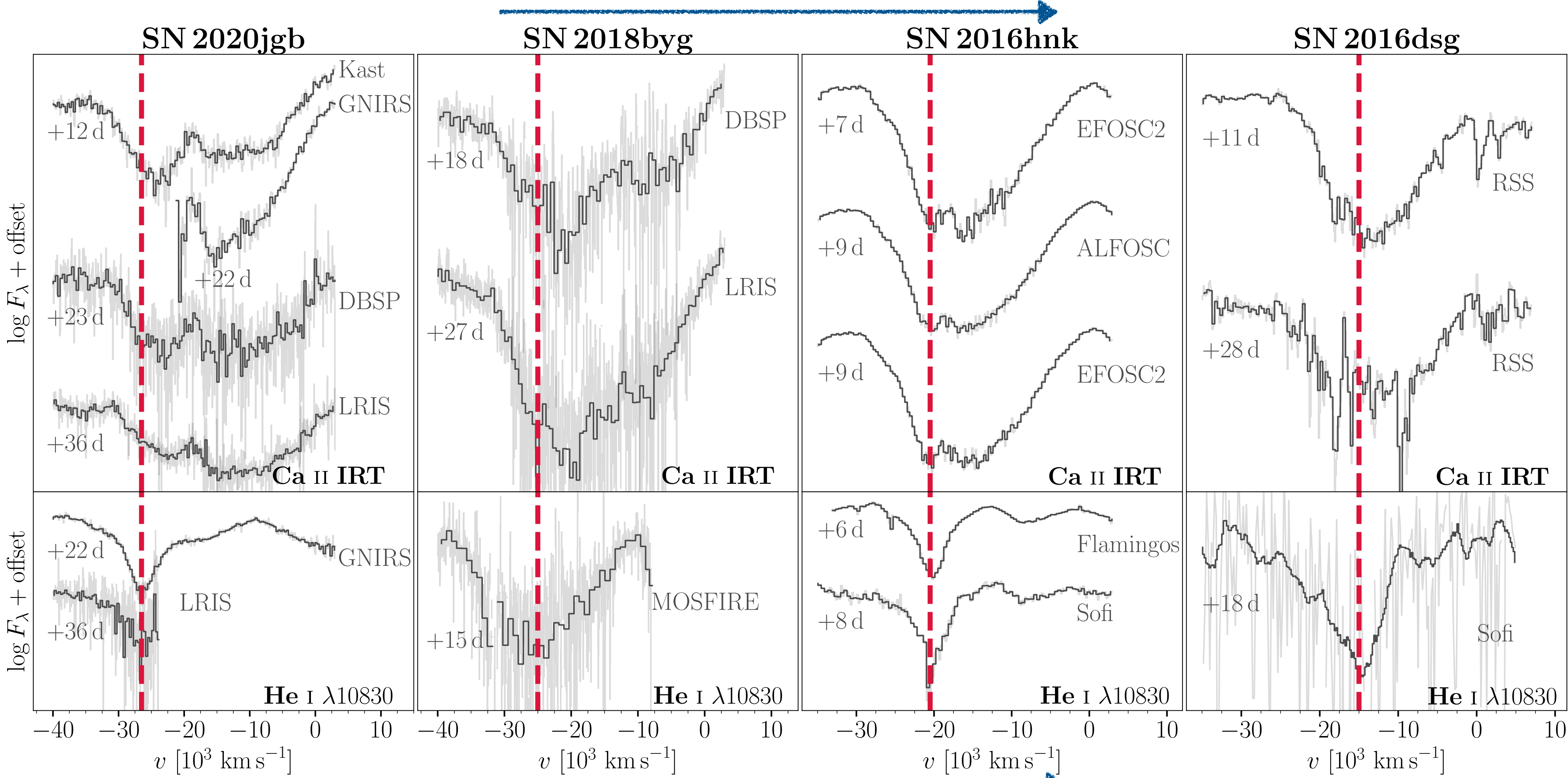
2D Model (Fink et al., 2010):
C/O core: 0.920 Msun
He shell: 0.084 Msun



Unburnt helium

He I $\lambda 10830$

Less prominent high-velocity features (HVF) of Ca II IRT



Lower "He I" velocity

Liu et al., 2023a

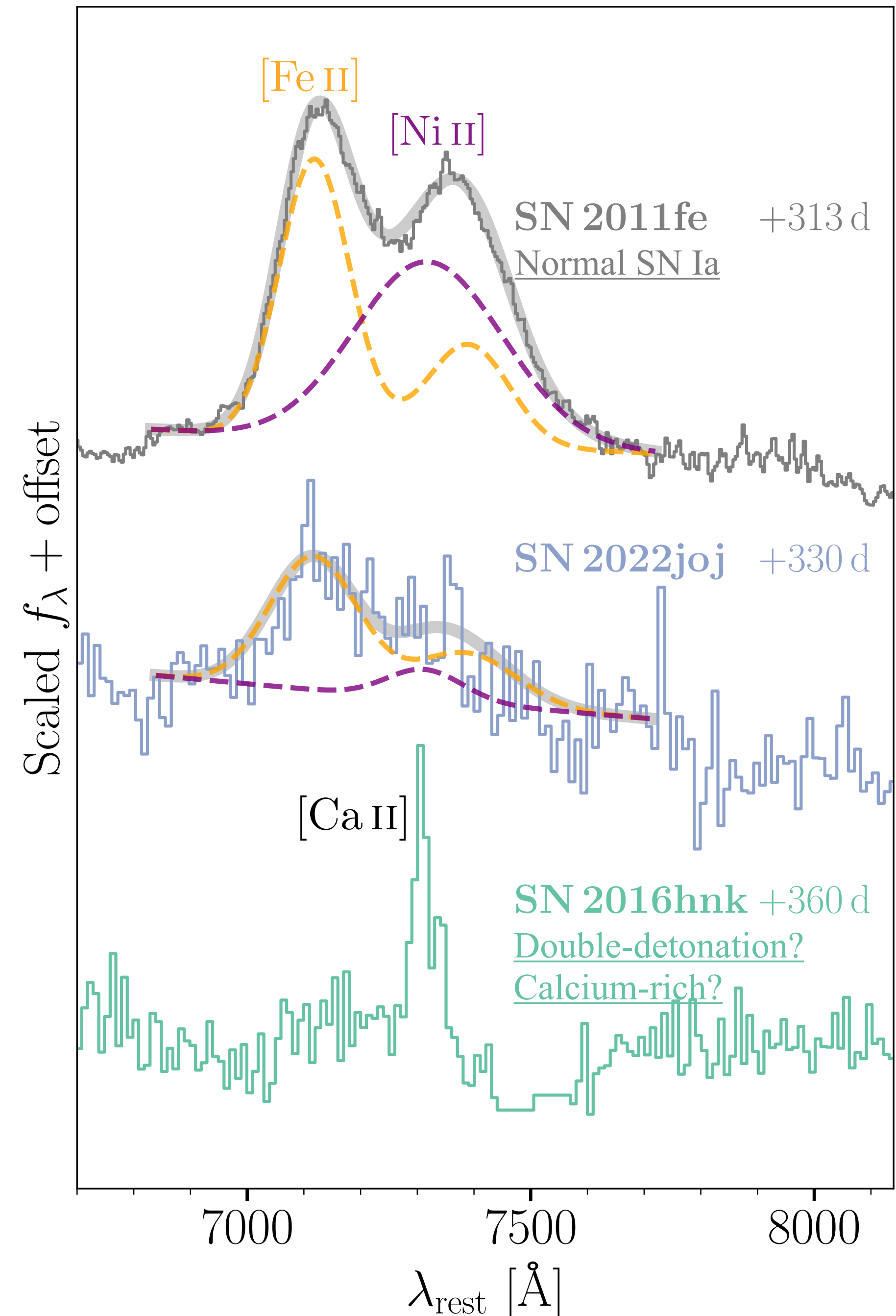
Nebular-phase Spectra

Ni/Fe Ratio

- **SN 2022joj** shows **little [Ni II]**
- Ni/Fe ratio < 0.03 (3σ)
- **Low Ni abundance** is consistent with a **sub- M_{Ch}** progenitor

Searching for [Ca II]

- Not detected in either SN 2020jgb or SN 2022joj



Uncovering Double Detonations

From the **first** to the **last** photon

- **Days after explosion**

- Flux excess — early discovery & photometry



La Silla Schmidt Southern Survey **(LS4)** is coming!

- **Around maximum luminosity**

- Dramatic color evolution — multi-band photometry
- Line blanketing — optical spectroscopy
- Unburnt helium — NIR spectroscopy

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- **Late time**

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